

# 4–12- and 25–34-GHz Cryogenic mHEMT MMIC Low-Noise Amplifiers

Beatriz Aja Abelán, *Member, IEEE*, Matthias Seelmann-Eggebert, Daniel Bruch, Arnulf Leuther, Hermann Massler, Boris Baldischweiler, Michael Schlechtweg, Juan Daniel Gallego-Puyol, *Member, IEEE*, Isaac López-Fernández, Carmen Diez-González, Inmaculada Malo-Gómez, Enrique Villa, and Eduardo Artal

**Abstract**—In this paper, monolithic microwave integrated circuit (MMIC) broadband low-noise amplifiers (LNAs) for cryogenic applications based on a 100-nm metamorphic high-electron mobility transistor (mHEMT) technology in combination with grounded coplanar waveguide are reported. A three-stage LNA, operating in 4–12 GHz and cooled to 15 K exhibits an associated gain of  $31.5 \text{ dB} \pm 1.8 \text{ dB}$  and average noise temperature of 5.3 K ( $\text{NF} = 0.079 \text{ dB}$ ) with a low power dissipation of 8 mW. Additionally another three-stage LNA 25–34 GHz cooled to 15 K has demonstrated a flat gain of  $24.2 \text{ dB} \pm 0.4 \text{ dB}$  with 15.2 K ( $\text{NF} = 0.22 \text{ dB}$ ), average noise temperature, with a very low power dissipation of 2.8 mW on chip. The mHEMT-based LNA MMICs have demonstrated excellent noise characteristics at cryogenic temperatures for their use in radio-astronomy applications.

**Index Terms**—Cryogenic low-noise amplifier (LNA), metamorphic high electron-mobility transistor (mHEMT), monolithic microwave integrated circuit (MMIC).

## I. INTRODUCTION

A NEW generation of focal plane arrays, with a large number of pixels, enhance the mapping efficiency of some of the existent radio telescopes. This need of a large number of receivers has been the initial motivation for the development of monolithic microwave integrated circuits (MMICs) especially designed to obtain good cryogenic performance. Cryogenic low-noise amplifiers (LNAs) working at 4–15-K ambient temperature are the key component in the front-ends of all those ultra-low-noise cryogenic receivers used for radio astronomy and deep-space communications. For such a large number of LNAs, MMICs enable simple,

small size, and low-cost production. These LNAs should be designed to obtain the lowest possible noise. Moreover, some of the radio-astronomy applications also require a very wide instantaneous bandwidth in order to increase the sensitivity of continuum observations. In that case, gain fluctuations of the amplifier may play a role since they can degrade the sensitivity.

To date, InP pseudomorphic high electron-mobility (pHEMT) MMIC LNAs have demonstrated outstanding noise at cryogenic temperatures [1]–[6]. However, recently metamorphic high electron-mobility transistor (mHEMT)-based LNAs have also been reported with good performances [7]–[10]. The advantages of mHEMT technology are lower costs, better robustness, and availability of larger GaAs wafers for production compared to InP substrate materials. The potential of the 100-nm mHEMT technology for MMIC applications at cryogenic temperatures is evaluated in this paper. We report the results obtained in two designs for the 4–12- and 25–34-GHz bands.

The 4–12-GHz band is of great interest since it is typically used in the IF of millimeter- and submillimeter-wave cryogenic receivers. The mixing element in this type of receiver is a superconducting–insulating–superconducting (SIS) junction, a superconducting hot-electron bolometer (HEB), or a Schottky-barrier diode, cryogenically cooled to 4 K.

The 25–34-GHz band is very important in very long baseline interferometry (VLBI) for astronomy and geodesy observation (see VLBI 2010 [11]), as well as for the down link of modern deep-space missions [12].

After a description of the technology in Section II, the device modeling with ambient and cryogenic measurements is shown in Section III. Circuit designs are discussed in Section IV, and their characterization at room and cryogenic temperatures is described in Section V.

## II. mHEMT TECHNOLOGY

For ultra-low-noise applications, high electron-mobility transistors (HEMTs) based on InGaAs/InAlAs heterostructures with high In-content in the electron transport channel are the most appropriate devices [13]. These heterostructures can be grown either directly on InP wafers or by using a metamorphic buffer to adapt the lattice constant on GaAs substrates. Major advantages of the metamorphic approach are cost and quality of the GaAs wafers. Furthermore, the material is less brittle compared to InP.

The InAlAs/InGaAs epi-structure is grown by molecular beam epitaxy (MBE) on 4-in semi-insulating GaAs substrates. For the metamorphic buffer, a 1- $\mu\text{m}$ -thick linear

Manuscript received July 10, 2012; revised September 13, 2012; accepted September 14, 2012. This work was supported by the Ministerio de Ciencia e Innovación under Research Program Grant TRA2009-0304. This work was supported in part by the Instituto Geográfico Nacional and the IAF-Fraunhofer funds. This paper is an expanded paper from the IEEE MTT-S International Microwave Symposium, Montreal, QC, Canada, June 17–22, 2012.

B. Aja Abelán, E. Villa, and E. Artal are with the Department of Communications Engineering, University of Cantabria, Santander 39005, Spain (e-mail: ajab@unican.es).

M. Seelmann-Eggebert, D. Bruch, A. Leuther, H. Massler, B. Baldischweiler, and M. Schlechtweg are with the Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg 79108, Germany (e-mail: Matthias.Seelmann-Eggebert@iaf.fraunhofer.de).

J. D. Gallego-Puyol, I. López-Fernández, C. Diez-González, and I. Malo-Gómez are with the Centro Astronómico de Yebes, Centro de Desarrollos Tecnológicos (CDT), Instituto Geográfico Nacional (IGN), Guadalajara 19080, Spain (e-mail: jd.gallego@oan.es).

Color versions of one or more of the figures in this paper are available online at <http://ieeexplore.ieee.org>.

Digital Object Identifier 10.1109/TMTT.2012.2221735

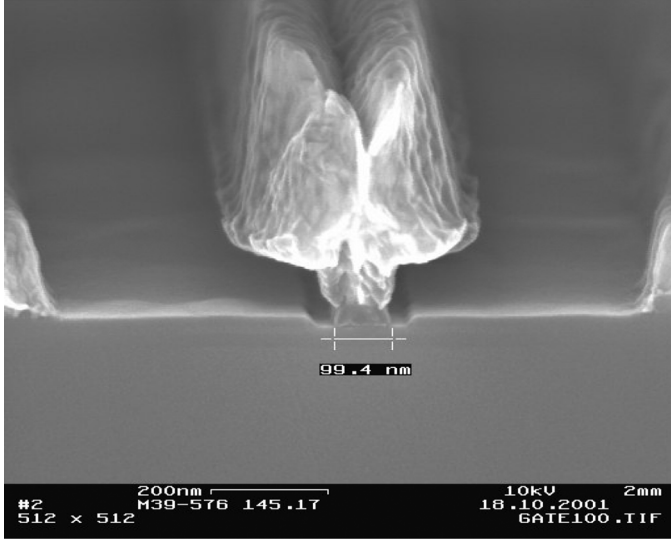


Fig. 1. SEM cross section of a 100-nm mHEMT.

$\text{In}_x\text{Al}_{0.48}\text{Ga}_{0.52-x}\text{As}$  ( $x = 0 \rightarrow 0.52$ ) transition is used. The electrons are confined in a  $\text{In}_{0.65}\text{Ga}_{0.35}\text{As}/\text{In}_{0.53}\text{Ga}_{0.47}\text{As}$  split channel to increase the breakdown voltage. The low-energy electrons are confined in the  $\text{In}_{0.65}\text{Ga}_{0.35}\text{As}$  layer with high electron mobility. In addition, the high-energy electrons are distributed over both layers, which reduces impact ionization. The split channel is confined by  $\text{In}_{0.52}\text{Al}_{0.48}\text{As}$  barriers. The upper barrier layer includes a silicon  $\delta$  doping. The layer sequence is capped with a highly doped  $\text{In}_{0.53}\text{Ga}_{0.47}\text{As}$  layer to reduce the ohmic contact and source resistance. A wet chemically mesa etch process is used for device isolation. The  $\text{InGaAs}$  channel layer is under-etched to avoid contact between the conducting  $\text{InGaAs}$  channel material and the gate metallization crossing the mesa edge in order to avoid gate leakage currents. Electron beam evaporated  $\text{GeAu}$  layers are used for the ohmic contacts, which are alloyed at 300 °C on a nitrogen purged hot plate. The 100-nm T gate is defined by 100-kV electron-beam lithography using a three-layer resist (PMMA). The gate recess is etched with a succinic-acid-based solution. Great care is taken to achieve homogeneous lateral etching across the 4-in wafer. An undersized lateral recess width would degrade the breakdown voltage of the device, whereas an oversized lateral recess would increase the source resistance [14]. A Pt Ti Pt–Au layer sequence is used for the gate metallization.

The active devices feature T-shaped 100-nm gates with an indium content of 65% in the main channel. A cross section of a realized 100-nm transistor is shown in Fig. 1. These devices also typically have a  $f_T$  of 220 GHz and a  $f_{\text{max}}$  of 300 GHz. A two-finger device with 0.06-mm total gate width is used for process monitoring. Typical values are a maximum transconductance of 1400 mS/mm with a drain bias of 1 V and a two-terminal reverse breakdown greater than 4 V at ambient temperature.

The transmission lines used in the MMICs are grounded coplanar waveguides with two metallization levels and 3- $\mu\text{m}$  Au thickness. The process further includes 50- $\Omega/\text{sq}$  NiCr thin-film resistors, 225 pF/mm<sup>2</sup> metal–insulator–metal (MIM) capacitors, via-holes, and CVD SiN passivation.

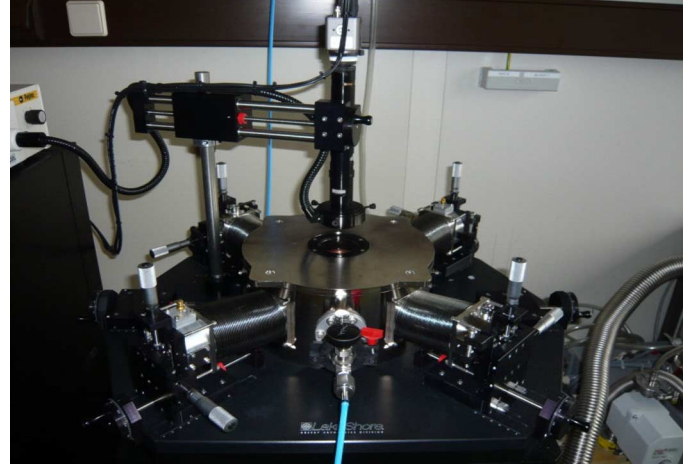


Fig. 2. Cryogenic probe station.

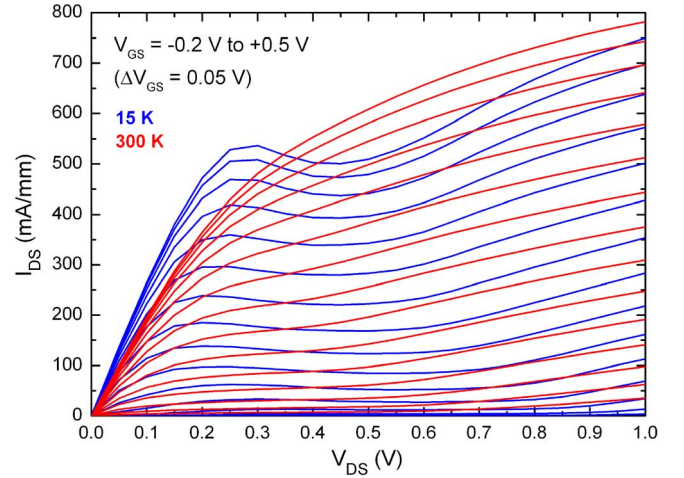


Fig. 3.  $I$ – $V$  characteristics for a  $4 \times 15 \mu\text{m}$  mHEMT device at 15 K (blue in online version) and at 300 K (red in online version) for several gate voltages.

The wafers are thinned down to a final thickness of 50  $\mu\text{m}$ . The 100-nm gate-length HEMT devices exhibited more than 90% yield for the fabricated wafers.

### III. DEVICE MODELING

The design of cryogenic LNAs requires a small-signal HEMT model with noise parameters; therefore, it was developed first [15]. The  $S$ -parameters of the device, both at ambient temperature 300 and 15 K, were measured on a wafer using a thru–reflect line (TRL) calibration, in the cryogenic probe station shown in Fig. 2. From these measurements, parasitic elements and intrinsic capacitances were determined. The noise parameters of the transistors were modeled following the approach of Pospieszalski [16]. Cryogenic models were improved based on the data from numerous runs.

DC measurements were carried out at 300 and 15 K on a representative device with gate length of 100 nm and gate width of  $4 \times 15 \mu\text{m}$ . The output  $I$ – $V$  characteristics are shown in Fig. 3. The device exhibits good pinch-off characteristics. The kink effect appears at around 15 K  $V_{\text{DS}} = 0.3$  V.

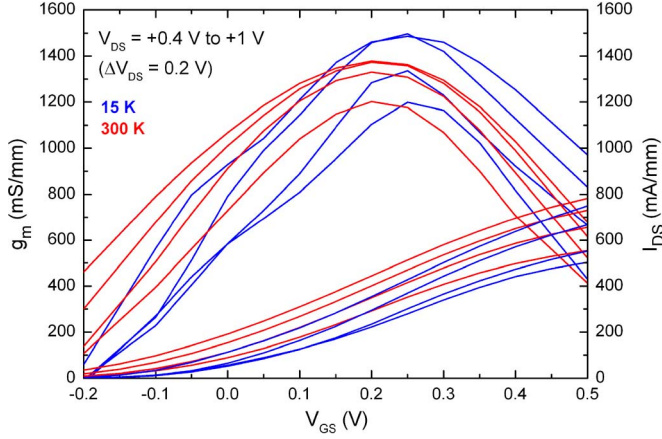


Fig. 4. Transconductance ( $g_m$ ) and  $I_{DS} - V_{GS}$  characteristics for a  $4 \times 15 \mu\text{m}$  mHEMT device at 15 and 300 K for several drain voltages.

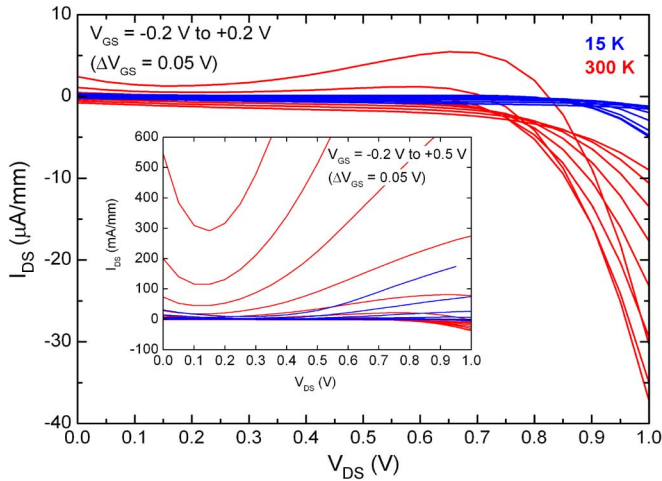


Fig. 5. Gate current-voltage characteristics for a  $4 \times 15 \mu\text{m}$  mHEMT device at 15 and at 300 K for several gate voltages.

The transconductance and  $I_{DS} - V_{GS}$  characteristics for the same device are plotted in Fig. 4. The device exhibits a maximum transconductance of 1500 mS/mm with  $V_{DS} = 0.6 \text{ V}$  at 15 K and 1400 mS/mm with  $V_{DS} = 1 \text{ V}$  at 300 K.

The gate leakage current is an important parameter related to the noise behavior of the device. Measured  $I_{GS} - V_{GS}$  characteristics are shown in Fig. 5. The gate leakage at 15 K is lower than at 300 K, in the range of a few nanoamperes for  $V_{GS}$  from  $-0.2$  to  $0.2 \text{ V}$  and  $V_{DS}$  from  $0.1$  to  $0.9 \text{ V}$ .

The device  $S$ -parameters were also measured on wafer at both ambient temperature 300 and 15 K up to 50 GHz. Measurement results can be found in [15] and [17].

#### IV. AMPLIFIERS DESIGN

To demonstrate the potential of the mHEMT technology at cryogenic temperatures, two MMIC LNAs were designed and manufactured. The two single-ended LNAs cover the frequency range from 4 to 12 GHz (IF for millimeter-wave receivers) and from 25- to 34-GHz VLBI and deep-space bands), respectively.

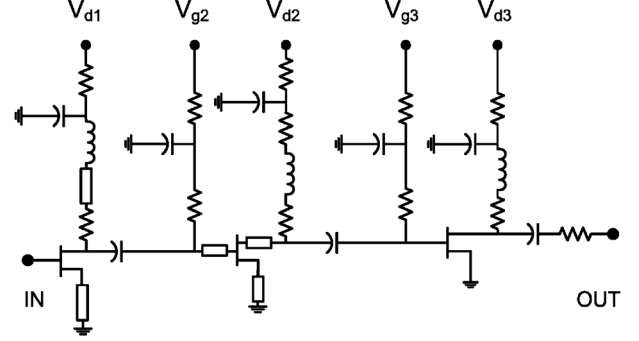


Fig. 6. Schematic diagram of the 4–12-GHz MMIC LNA.

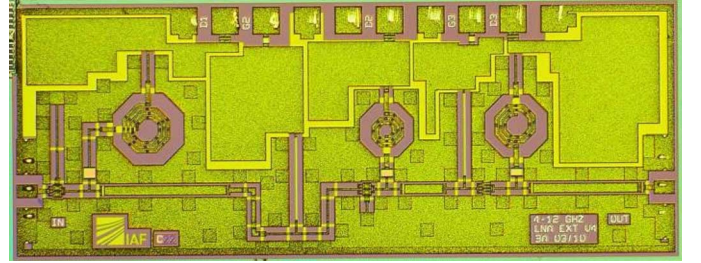


Fig. 7. Photograph of the manufactured 4–12-GHz MMIC LNA. The chip size is  $2.5 \text{ mm} \times 1 \text{ mm}$ .

#### A. LNA 4–12 GHz

A simplified schematic of the three-stage 4–12-GHz LNA is shown in Fig. 6. The LNA has three stages, which include  $4 \times 30 \mu\text{m}$  transistors. A photograph of the three-stage 4–12-GHz LNA is shown in Fig. 7. The chip size is  $2.5 \text{ mm} \times 1 \text{ mm}$ .

The first stage is mainly optimized for minimum noise figure, while the second stage is matched partially for noise, and the third stage fully for gain. Gain flatness and input and output reflection coefficients were also taken into account during the design optimization. Grounded coplanar-waveguide lines are used for matching networks and interconnections. Bias is brought to the transistors through resistors and stubs capacitively shorted with RF bypass MIM capacitors. The resistors act as damping elements suppressing the excess out-of-band gain and stabilizing the amplifier.

Matching the first transistor for minimum noise requires a high impedance of about  $100 \Omega$  real part with low losses. An external matching network, on a low-loss substrate, allows more flexibility to match the optimum noise impedance to the source impedance at these frequencies. Therefore, an external microstrip input matching network has been designed as shown Fig. 8, which is fabricated on a 0.254-mm CLTE-XT Arlon substrate with lower losses than GaAs, and the inductance is realized by one 1.6-mm-long bond wire.

This network allows to achieve broadband matching, to reduce the circuit noise figure, and to tune the amplifier. The first gate is biased through the RF input, and this is done through the external input matching network.

In order to characterize its performance, the MMIC was assembled in a gold-plated brass module with SMA coaxial connectors, shown in Fig. 9.



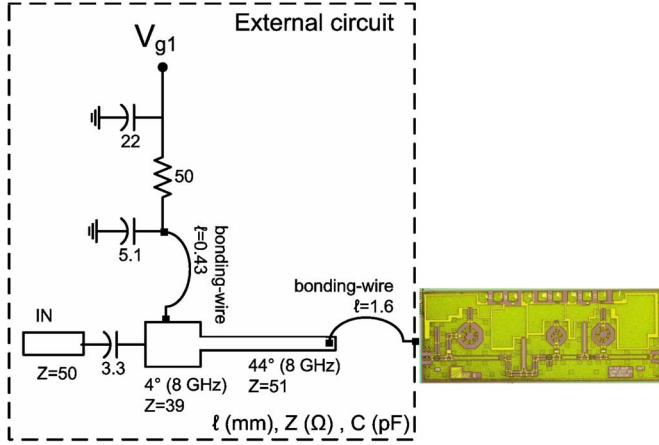


Fig. 8. Schematic diagram of a 4–12-GHz MMIC LNA with the external input matching network.

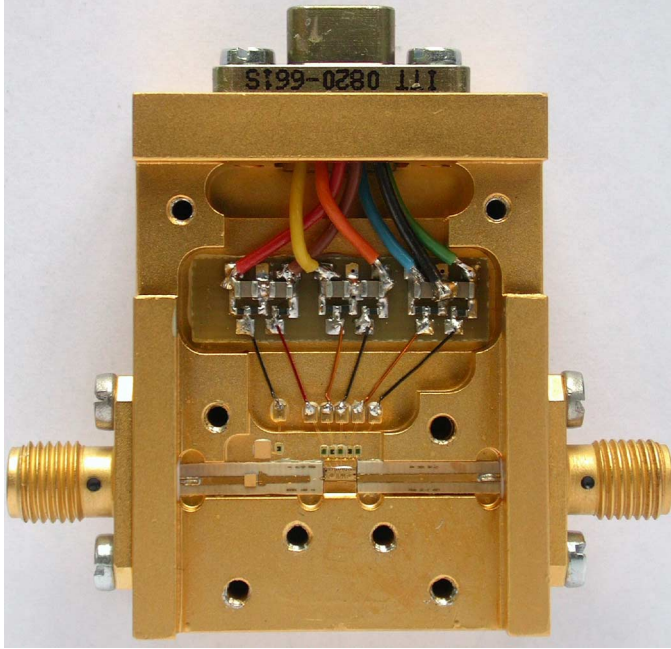


Fig. 9. Packaged module of the 4–12-GHz LNA. The box dimensions are 31.3 mm  $\times$  40 mm  $\times$  15 mm.

### B. LNA 25–34 GHz

As is shown in Fig. 10, the MMIC LNA consists of three stages; each one employs a  $4 \times 15 \mu\text{m}$  transistor, with a source inductive stub in the two first stages. Grounded coplanar-waveguide lines are used for matching networks and interconnection. To accomplish wideband matching, lossy matching networks are used both for interstage matching and for biasing the drains.

A photograph of the three-stage LNA is shown in Fig. 11. The chip size is 2.5 mm  $\times$  1 mm.

The LNA MMIC was assembled in a gold-plated aluminium module with 2.92-mm coaxial connectors, shown in Fig. 12. Gold-plated microstrip lines on a 0.127-mm CLTE-XT Arlon dielectric substrate were used at the input and output of the MMIC.

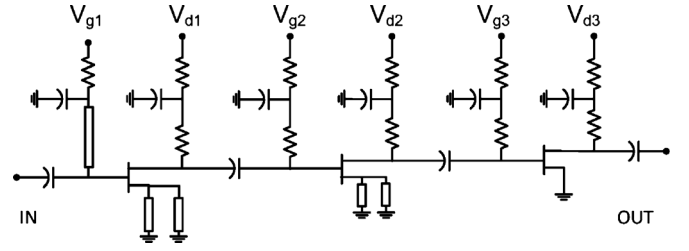


Fig. 10. Schematic diagram of the 25–34-GHz MMIC LNA.

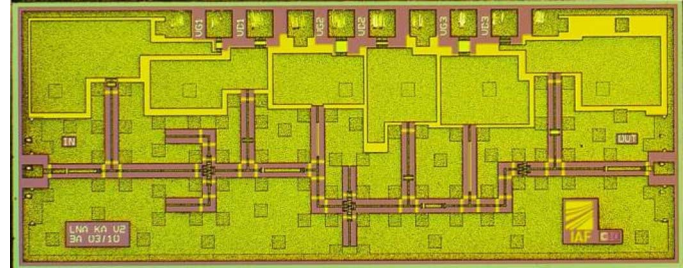


Fig. 11. Photograph of the manufactured 25–34-GHz MMIC LNA. The chip size is 2.5 mm  $\times$  1 mm.

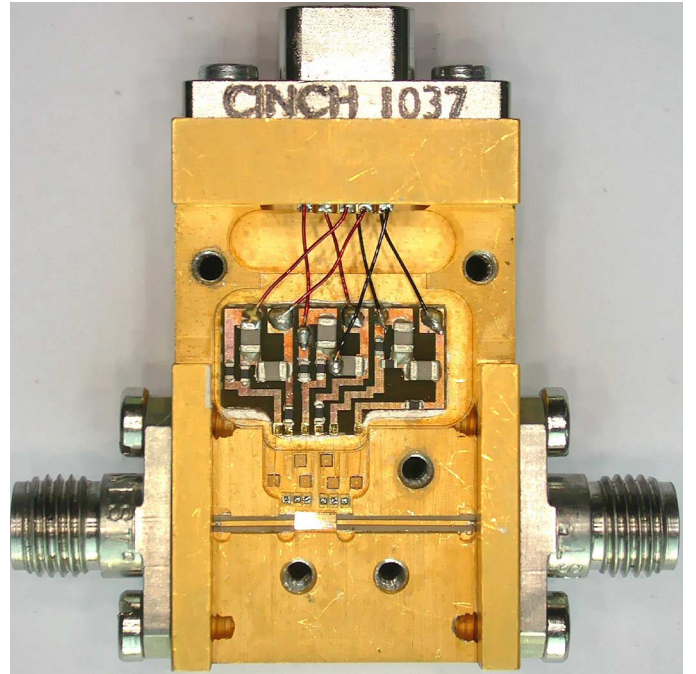


Fig. 12. Packaged module of the 25–34-GHz LNA. The dimensions of the box are 20.6 mm  $\times$  32.5 mm  $\times$  12 mm.

## V. EXPERIMENTAL RESULTS

$S$ -parameters, noise temperature, and gain fluctuations were tested at 300- and 15-K temperatures. The chips were first tested on-wafer and then the packaged LNAs were measured at 300 and 15 K. The characterization at cryogenic temperatures of the modules was carried out in a closed-cycle helium cryostat, partially shown in Fig. 13.

For both amplifiers, two units were packaged and tested at cryogenic temperatures in order to verify the measurements. Test results demonstrated very good repeatability, and therefore that the process is sufficiently stable. Moreover, the amplifiers

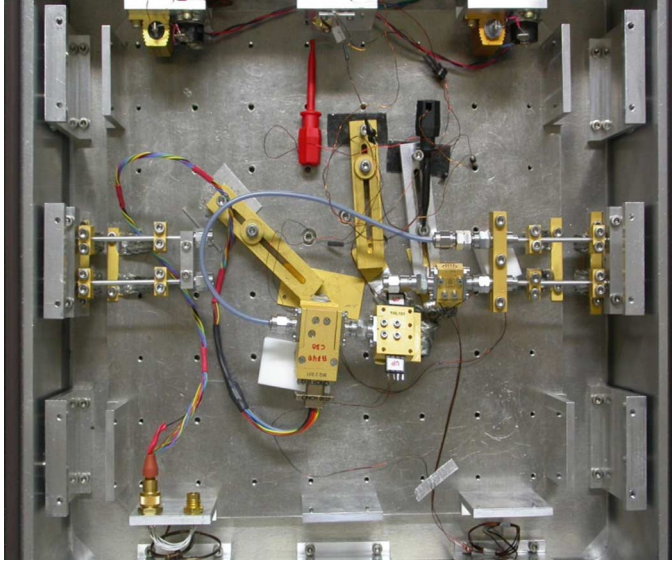


Fig. 13. Amplifier inside the vacuum chamber of the cryostat test system.

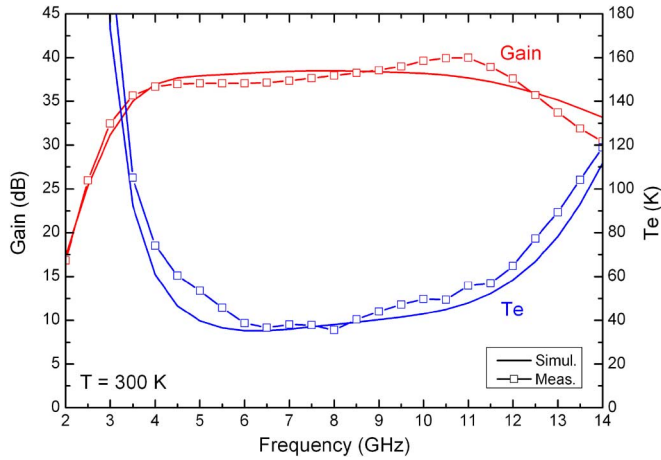


Fig. 14. Measured and simulated gain and noise temperature of the LNA 4–12 GHz at 300 K.

were unconditionally stable at both room and cryogenic temperatures, which was checked from the  $S$ -parameter measurement.

#### A. LNA 4–12 GHz

The 4–12-GHz amplifier was tested at room temperature and it exhibited the gain and noise temperature depicted in Fig. 14. The measured gain was 38 dB with ripple of 1.6 dB in the 4–12-GHz frequency band and with average noise of 47.8 K. The on-chip power consumption for this optimum noise bias was 90.2 mW;  $V_d = 1.5$  V,  $I_d = 60$  mA (166 mA/mm). The minimum noise was 35 K with an associated gain of 37.7 dB at 7.75 GHz.

At 15 K, the cold attenuator method was used to measure the LNA module noise performance, where the incoming noise power is generated by a commercial noise source. A cooled 15-dB attenuator provides a well-defined cold-temperature noise reference at the input of the LNA and reduces the change of reflection coefficient between the on- and off-states of the noise source. In this way, the measurement uncertainty of the noise temperature is reduced to  $\pm 1.4$  K.

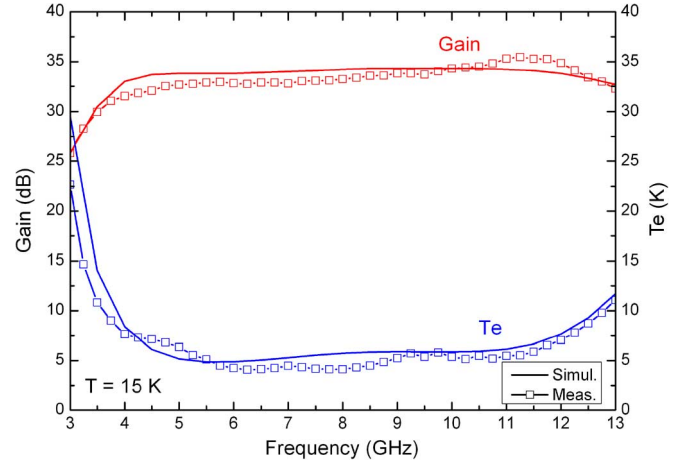
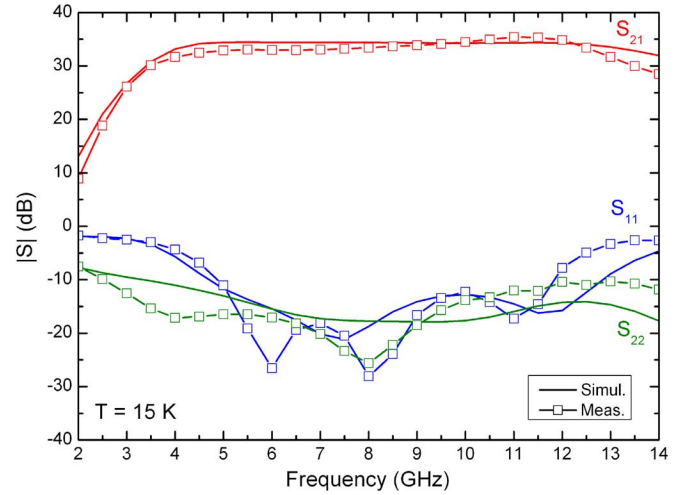


Fig. 15. Measured and simulated gain and noise temperature of the LNA 4–12 GHz at 15 K.

Fig. 16. Measured and simulated  $S$ -parameters of the LNA 4–12 GHz at 15 K.

The measured and simulated gain and noise temperature at 15 K from 3 to 13 GHz are shown in Fig. 15. The simulation comprises the MMIC LNA, as well as the external microstrip network and coaxial connectors. From 4 to 12 GHz, the amplifier achieves 31.5-dB gain with  $\pm 1.8$ -dB flatness. The average noise temperature is 5.3 K from 4 to 12 GHz. The measurement was made with a drain voltage  $V_d = 0.53$  V and a total drain current of 15 mA (41.6 mA/mm). The dc power dissipation is 8 mW on-chip. The minimum noise temperature is 4 K at 6.25 GHz with 30 dB of gain.

The LNA  $S$ -parameters at cryogenic temperature (15 K) were also measured referred to the input and output ports of the packaged amplifier for the same bias conditions as in the measurement achieving the best noise performance, shown in Fig. 16. Input and output return losses are greater than 5 and 14 dB in the band, respectively (the input reflection loss (IRL) is greater than 12 dB from 5 to 12 GHz).

#### B. LNA 25–34 GHz

After testing the  $S$ -parameters of the chip on-wafer at room temperature, it was also tested in the cryogenic probes station at 15 K. The  $S$ -parameters at 15 K are shown in Fig. 17. With



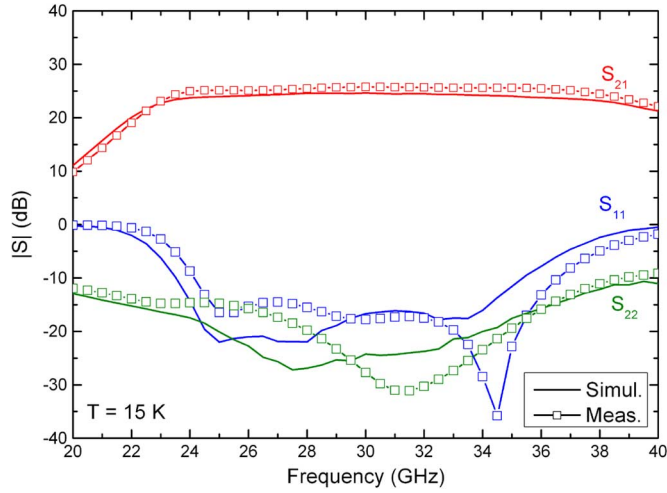
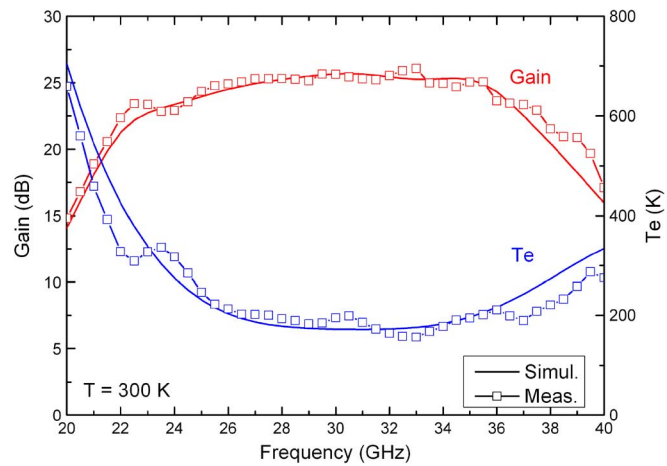
Fig. 17.  $S$ -parameters on-wafer of the LNA 25–34 GHz at 15 K.

Fig. 18. Measured and simulated gain and noise temperature of the packaged LNA 25–34 GHz at 300 K.

a power dissipation of 3.6 mW,  $V_d = 0.4$  V and a total current of 9.0 mA (50 mA/mm), the gain was 24.3 dB with a ripple of only 0.3 dB. Input and output return losses were greater than 15 and 20 dB in the band, respectively.

The amplifier was then measured in the module with coaxial connectors at 300 K. The gain and noise temperature are depicted in Fig. 18.

The measured gain was 25.2 dB with ripple of 0.8 dB in the 25–34-GHz frequency band. The average noise in the band was 190 K with a power dissipation of 17.3 mW on chip;  $V_d = 0.68$  V,  $I_d = 25.5$  mA (141 mA/mm). The minimum noise temperature was 183 K at 29 GHz, with an associated gain of 25 dB.

Next, the noise performance was measured with the amplifier cooled to cryogenic temperature. The measured and simulated gain and noise temperature from 20 to 40 GHz at 15 K are shown in Fig. 19. The simulation includes the amplifier with bonding wires, microstrip lines, and coaxial connectors. The gain and noise were tested using the cold attenuator method. The measured gain was 24.2 dB with a ripple of 0.4 dB in the 25–34-GHz frequency band and the average noise in the band was 15.2 K.

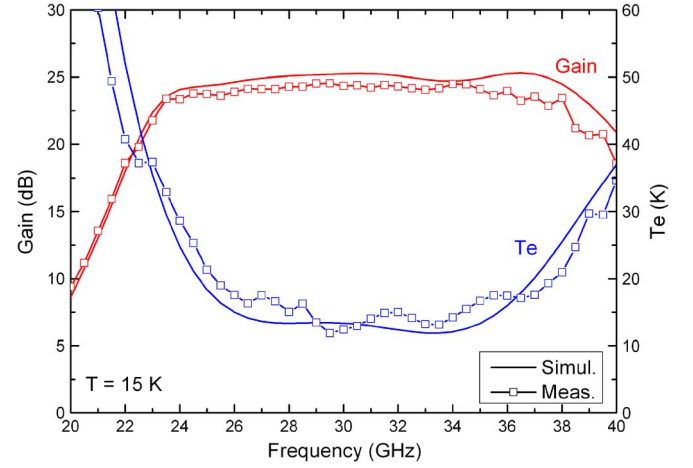
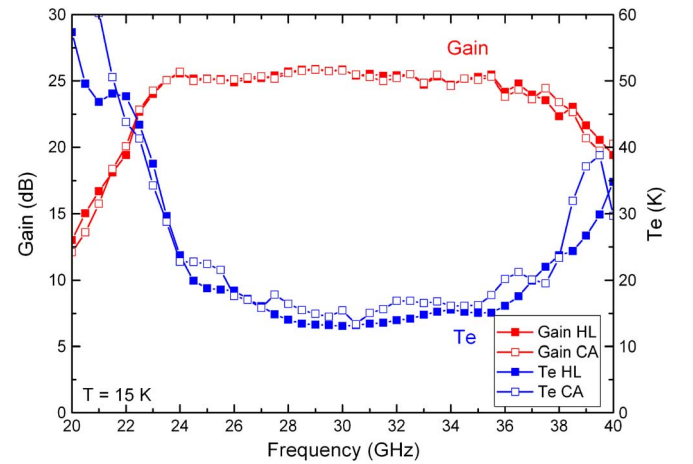


Fig. 19. Measured and simulated gain and noise temperature of the packaged LNA 25–34 GHz at 15 K.

Fig. 20. Measured gain and noise temperature of the packaged LNA 25–34 GHz at 15 K, using the cold attenuator (CA) method with a diode as noise source, and the hot/cold method (HL) with a 50- $\Omega$  load at the input of the amplifier.

The minimum noise temperature was 11.8 K with an associated gain of 24.5 dB at 29.5 GHz. The power consumption was 2.8 mW on chip;  $V_d = 0.37$  V and  $I_d = 7.5$  mA (42 mA/mm).

An identical LNA from the same wafer was also assembled and tested at cryogenic temperature using two different methods in order to verify the measurements of noise and gain. The methods were: 1) the cold attenuator method with a diode noise source and 2) the hot/cold method with a 50- $\Omega$  load at the input of the amplifier [18]. The results obtained in gain and noise temperature for both methods are shown in Fig. 20. The gain was the same,  $25.3 \pm 0.6$  dB and the average temperature 16.6 K with the cold-attenuator method, and 14.9 K with the hot/cold method, showing a difference in the average temperature of only 1.7 K. In addition, the hot/cold method showed a smoother response of the noise temperature versus frequency.

One of the possible applications of this amplifier is in the front end of cosmic microwave background (CMB) radiation receivers [19], [20]. The very wide band of those receivers combined with the extreme low noise makes the problem of gain fluctuations more prominent. The gain fluctuation of the LNA

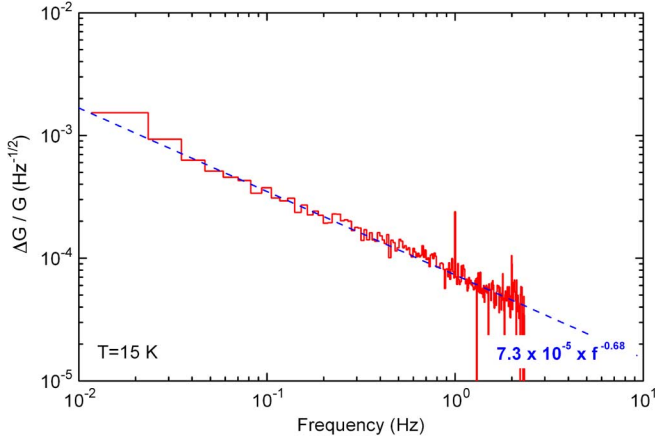


Fig. 21. Normalized gain fluctuation spectrum of the LNA 25–34 GHz at 15 K. The dashed line indicates the noise floor of the measurement system.

TABLE I  
COMPARISON OF WIDE-BANDWIDTH LNA AT CRYOGENIC TEMPERATURE

Ref.	Process	Type	Freq (GHz)	Gain (dB)	Te @ Tamb (K)	Pdis (mW)
[1]	0.1 μm InP HEMT	3-stages MMIC	1-11	33.4±0.3	3.9 @ 4.1	24
[2]	0.1 μm InP HEMT	3-stages MMIC	4-12	37	3.9 @ 12	9.2
[6]	0.13 μm InP HEMT	3-stages MMIC	0.5-13	38	4.4 @ 12	16.5
[8]	0.1 μm mHEMT	3-stages MMIC	4-12	26±1.2	8.1 @ 15	12
This Work	0.1 μm mHEMT	3-stages MMIC	4-12	31.5±1.8	5.3 @ 15	8
[3]	0.1 μm InP HEMT	3-stages MMIC	26-40	23±1.1	15.5 @ 12	5.95
		2 cascaded 3-stages MMIC	26-40	41±2.4	11.4 @ 12	--
[4]	InP HEMT	Hybrid	27-37(*)	28±7	14.5 @ 12	--
[5]	InP HEMT	4-stages MMIC	26-36	42	20.0 @ 15	--
This Work	0.1 μm mHEMT	3-stages MMIC	25-34	24.2±0.4	15.2 @ 15	2.8

(\*) Narrowband design with  $T_e = 9$  K and Gain = 30 dB at 32 GHz when cooled at 12 K.

25–34 GHz operating cryogenically was measured as described in [21]. The spectrum of normalized gain fluctuation (SNGF) measured is shown in Fig. 21. The SNGF is usually modeled by an expression of the form [21]

$$S(f) = \beta \cdot \left( \frac{1 \text{ Hz}}{f} \right)^\alpha \quad (1)$$

where  $S(f)$  is the unilateral spectral density (i.e., SNGF) with units  $\text{Hz}^{-1/2}$ . In cryogenic LNAs, the parameter  $\alpha$  is usually close to 0.5. The best fit to the data shown in Fig. 21 is obtained with  $\beta = 7.3 \cdot 10^{-5} \text{ Hz}^{-(1/2)}$  and  $\alpha = 0.674$ .

These results would meet the stringent specifications of Atacama Large Millimeter/Sub-millimeter Array (ALMA)<sup>1</sup> or

<sup>1</sup>ALMA, Santiago, Chile. [Online]. Available: <http://www.almaobservatory.org/>

HERSCHEL [22], achieving SNFG values at 1 Hz lower than those obtained for cryogenic InP amplifiers [23].

A comparison of the two LNAs with other reported wideband MMIC LNAs, working in the same frequency ranges at cryogenic temperature, is given in Table I. Almost all data found in the literature refer to InP-based devices. Both presented LNAs offer very low-noise performance with ultra-low-power dissipation at cryogenic temperatures.

## VI. CONCLUSION

Ultra-low-noise MMIC LNAs that can operate cryogenically cooled are required for radio-astronomy applications. We reported the fabrication and characterization of two MMIC LNAs based on 100-nm mHEMT technology. These very low-noise and wide instantaneous bandwidth amplifiers have been developed in order to demonstrate the excellent performance of this mHEMT technology at cryogenic temperatures. Their performance has been characterized in terms of small-signal gain and equivalent noise temperature under cryogenic operating conditions (15 K). A three-stage LNA demonstrated a small-signal gain of 31.5 dB and average noise temperature of 5.3 K from 4 to 12 GHz when cooled at 15 K with only 8.0-mW power consumption.

The 25–34-GHz MMIC LNA exhibited a gain of 24.2 dB and average noise temperature of 15.2 K with 2.8-mW power consumption. Both LNAs offer very low-noise performance with very low-power dissipation at cryogenic temperatures. The presented results demonstrate the high potential of mHEMT technology for cryogenically cooled very sensitive wideband receivers.

## ACKNOWLEDGMENT

The authors would like to thank all the people involved in manufacturing the prototypes used in this study. The authors express their gratitude to the Technology Department, Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany, for the epitaxial growth and wafer processing.

## REFERENCES

- [1] J. Randa, E. Gerecht, D. Gu, and R. L. Billinger, "Precision measurement method for cryogenic amplifier noise temperatures below 5 K," *IEEE Trans. Microw. Theory Techn.*, vol. 54, no. 3, pp. 1180–1189, Mar. 2006.
- [2] J. D. Pandian, L. Baker, G. Cortes, P. F. Goldsmith, A. A. Deshpande, R. Ganesan, J. Hagen, L. Locke, N. Wadefalk, and S. Weinreb, "Low noise 6–8 GHz receiver," *IEEE Microw. Mag.*, vol. 7, no. 6, pp. 74–84, Dec. 2006.
- [3] Y.-L. Tang, N. Wadefalk, M. A. Morgan, and S. Weinreb, "Full Ka-band high performance InP MMIC LNA module," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2006, pp. 81–84.
- [4] J. J. Bautista, J. G. Bowen, N. E. Fernandez, Z. Fujiwara, J. Loreman, S. Petty, S. J. L. Prater, R. Grunbacher, R. Lai, M. Nishimoto, M. R. Murti, and J. Laskar, "Cryogenic, X-band and Ka-band InP HEMT based LNAs for the deep space network," in *IEEE Aerosp. Conf.*, Mar. 2001, vol. 2, pp. 829–842.
- [5] D. Kettle, N. Roddis, and R. Sloan, "A lattice matched InP chip set for a Ka band radiometer," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2005, pp. 1033–1036.
- [6] J. Schlee, N. Wadefalk, P. Nilsson, P. Starski, G. Alestig, J. Halonen, B. Nilsson, A. Malmros, H. Zirath, and J. Grahn, "Cryogenic 0.5–13 GHz low noise amplifier with 3 K midband noise temperature," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2012, pp. 1–3.

- [7] A. Leuther, A. Tessmann, I. Kallfass, R. Losch, M. Seelmann-Eggebert, N. Wadefalk, F. Schafer, J. D. G. Puyol, M. Schlechtweg, M. Mikulla, and O. Ambacher, "Metamorphic HEMT technology for low-noise applications," in *IEEE Int. Indium Phosphide Relat. Mater. Conf.*, 2009, pp. 188–191.
- [8] B. Aja, K. Schuster, F. Schafer, J. D. Gallego, S. Chartier, M. Seelmann-Eggebert, I. Kallfass, A. Leuther, H. Massler, M. Schlechtweg, C. Diez, I. Lopez-Fernandez, S. Lenz, and S. Turk, "Cryogenic low-noise mHEMT-based MMIC amplifiers for 4–12 GHz band," *IEEE Microw. Wireless Compon. Lett.*, vol. 21, no. 11, pp. 613–615, Nov. 2011.
- [9] S.-H. Weng, H.-Y. Chang, C.-C. Chiong, and M.-T. Chang, "Cryogenic evaluation of a 30–50 GHz 0.15- $\mu$ m MHEMT low noise amplifiers for radio astronomy applications," in *Proc. 41st Eur. Microw. Conf.*, Oct. 2011, pp. 934–936.
- [10] B. Aja, M. Seelmann-Eggebert, A. Leuther, H. Massler, M. Schlechtweg, C. Diez, J. D. Gallego, I. Lopez-Fernandez, I. Malo, E. Artal, and E. Villa, "4–12 GHz and 25–34 GHz cryogenic MHEMT MMIC low noise amplifiers for radio astronomy," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2012, pp. 1–3.
- [11] A. Niell, A. Whitney, B. Petrachenko, W. Schlüter, N. Vandenberg, H. Hase, Y. Koyama, C. Ma, H. Schuh, and G. Tuccari, "VLBI2010: Current and future requirements of geodetic VLBI systems," NASA, Washington, DC, Int. VLBI Service for Geodesy and Astrometry 2005 Annu. Rep. NASA/TP-2006-214136, 2006, pp. 13–40. [Online]. Available: <ftp://ivsc.gsfc.nasa.gov/pub/annual-report/2005/pdf/spcl-vlbi2010.pdf>, D. Behrend and K. Baver, Eds.
- [12] E. Vassallo, R. Martin, R. Madde, M. Lanucara, P. Besso, P. Droll, G. Galtie, and J. De Vicente, "The European Space Agency's deep-space antennas," *Proc. IEEE*, vol. 95, no. 11, pp. 2111–2131, Nov. 2007.
- [13] L. Samoska, "Towards terahertz MMIC amplifiers: Present status and trends," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2006, pp. 333–336.
- [14] K. Shinohara, Y. Yamashita, A. Endoh, K. Hikosaka, T. Matsui, S. Hiyamizu, and T. Mimura, "Importance of gate-recess structure to the cutoff frequency of ultra-high-speed InGaAs/InAlAs HEMTs," in *14th Indium Phosphide Relat. Mater.*, 2002, pp. 451–454.
- [15] M. Seelmann-Eggebert, F. Schaefer, A. Leuther, and H. Massler, "A versatile and cryogenic mHEMT-model including noise," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2010, pp. 501–504.
- [16] M. W. Pospieszalski, "Modeling of noise parameters of MESFET's and MODFET's and their frequency and temperature dependence," *IEEE Trans. Microw. Theory Techn.*, vol. 37, no. 9, pp. 1340–1350, Sep. 1989.
- [17] M. Seelmann-Eggebert, A. Leuther, H. Maßler, B. Aja, D. Bruch, A. Tessmann, I. Kallfass, and M. Schlechtweg, "The IAF mHEMT low-noise technology and its extension to cryogenic applications," presented at the 1st Eng. Forum Workshop, Göteborg, Sweden, Jun. 23–24, 2009.
- [18] D. Bruch, F. Schafer, B. Aja, A. Leuther, M. Seelmann-Eggebert, I. Kallfass, M. Schlechtweg, and O. Ambacher, "A single chip broadband noise source for noise measurements at cryogenic temperatures," in *IEEE MTT-S Int. Microw. Symp. Dig.*, Jun. 2011, pp. 1–4.
- [19] M. Bersanelli *et al.*, "Planck pre-launch status: Design and description of the low frequency instrument," *Astron. Astrophys.*, vol. 520, Oct. 2010, Art. ID AN. A4.
- [20] J. A. Rubiño-Martin *et al.*, "The QUIJOTE CMB experiment," in *Highlights of Spanish Astrophysics V, Astrophysics and Space Science Proceedings*. Berlin, Germany: Springer, 2010, pt. 3, pp. 127–135.
- [21] J. D. Gallego, I. López-Fernández, C. Diez, and A. Barcia, "Methods for the characterization and measurement of the gain fluctuations of cryogenic amplifiers," ALMA, Santiago, Chile, ALMA Memo 560, Oct. 2006. [Online]. Available: <http://www.alma.nrao.edu/memos/html-memos/alma560/memo560.pdf>.
- [22] T. de Graauw *et al.*, "The Herschel-Heterodyne Instrument for the far-infrared (HIFI)," *Astron. Astrophys.*, vol. 518, L6, pp. 1–7, Jul.–Aug. 2010.
- [23] I. López-Fernández, J. D. Gallego, C. Diez, A. Barcia, and J. M. Pintado, "Wide band, ultra low noise cryogenic InP IF amplifiers for the Herschel mission radiometers," *Proc. SPIE*, vol. 4855, pp. 489–500, 2003.



**Beatriz Aja Abelán** (S'01–A'05–M'06) received the Telecommunications Engineering degree and Ph.D. degree from the University of Cantabria, Santander, Spain, in 1999 and 2007 respectively.

Since 1999, she has been with the Department of Communications Engineering, University of Cantabria. From 2008 to 2012, she was an Invited Scientist with the Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany, in a joint collaboration project with Centro Astronómico de Yebes (CAY). Her research interests are in the area of microwave and millimeter-wave circuits and systems design, in particular the design of microwave LNAs for cryogenic applications.



**Matthias Seelmann-Eggebert** received the Diploma and Ph.D. degree in physics from the University of Tübingen, Tübingen, Germany, in 1980 and 1986, respectively.

From 1980 to 1996, he was involved in research and development related to infrared detectors based on HgCdTe and applied research in the field of electrochemistry and surface physics. From 1990 to 1991, he was a Visiting Scientist with Stanford University, Stanford, CA. From 1997 to 2000, he was engaged in the growth of CVD diamonds. Since 2001, he has been a member of the High Frequency Devices and Circuits Department, Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany, where he is in charge of the development of linear and nonlinear compact models for III–V devices.



**Daniel Bruch** was born in 1982. He received the Dipl.-Ing. degree from the University of Karlsruhe, Karlsruhe, Germany, in 2008, and is currently working toward the Ph.D. degree at the Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany.

His main research interests are the design and characterization of cryogenic LNAs.



**Arnulf Leuther** received the Dipl. Phys. degree and Ph.D. degree in physics from the Technical University of Aachen, Aachen, Germany.

From 1992 to 1996, he was with Forschungszentrum Jülich. In 1996, he joined the Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany, where he is currently Head of the Lithography Group. His research is focused on the development of advanced III–V process technologies for metamorphic HEMTs and the fabrication of millimeter- and submillimeter-wave

MMICs.

Dr. Leuther received the 1996 Borchert Medal from RWTH Aachen for his dissertation thesis.



**Hermann Massler** was born in Radolfzell, Germany, in 1965. He studied electrical engineering at the Technical University Karlsruhe, Karlsruhe, Germany, where he graduated in 1993.

While working on the Diploma degree at the Forschungszentrum Karlsruhe (FZK), he performed and investigated quasi-optical measurements at 140 GHz. He continued these studies as an FZK Research Assistant for an additional year. Since 1994, he has been with the Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany, where he



is involved with transistor and integrated circuit (IC) characterization up to 325 GHz.



**Boris Baldischweiler** was born in Freiburg, Germany, in 1984. He received the Dipl. Ing. (FH) degree in electrical engineering and M.Sc. degree in microsystems engineering from Hochschule Furtwangen University, Furtwangen, Germany, in 2009 and 2010, respectively.

Since 2010, he has been with the Department High Frequency Devices and Circuits, Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany. His main research concerns noise and cryogenic measurements.



**Michael Schlechtweg** received the Dipl.-Ing. Degree in electrical engineering from the Technical University Darmstadt, Darmstadt, Germany, in 1982, and the Dr.-Ing. degree from the University of Kassel, Kassel, Germany, in 1989.

He then joined the Fraunhofer Institute for Applied Solid State Physics (IAF), Freiburg, Germany, where he has been involved with the design of millimeter-wave integrated circuits and nonlinear characterization and modeling of active RF devices. In 1994, he became Head of the Simulation and Modeling Group,

Fraunhofer IAF. Since 1996, he has led the RF Devices and Circuits Department, where he has focused on the design and characterization of devices and integrated circuits based on III–V compound semiconductors for RF applications, as well as the development of integrated circuits and modules for sensor and communication systems up to 500 GHz and above. He coauthored approximately 200 scientific publications. He holds two patents.

Dr. Schlechtweg was the recipient of the 1993 Fraunhofer Prize and the 1998 European Microwave Prize.



**Juan Daniel Gallego-Puyol** (M'91) was born in Madrid, Spain, in 1960. He received the Doctor degree in physics in 1992.

Since 1985, he has been with the Observatorio Astronómico Nacional (OAN). In 1989, he spent one year with the National Radio Astronomy Observatory. His main research activity has been the development of cryogenic LNAs. He has been involved in numerous international projects in this field. Among others, he has been in charge of the development and construction of cryogenic amplifiers of the Herschel

European Space Agency (ESA) mission and of amplifiers for the European contribution to ALMA.

Dr. Gallego is a member of URSI and IAU.



**Isaac López-Fernández** was born in Oviedo, Spain, in 1969. He received the Eng. degree in telecommunication in 1995.

In 1994, he joined the Observatorio Astronómico Nacional (OAN), where he was initially involved with the development of receivers for VLBI. In 1995, he was involved with VLBI research with the Harvard Smithsonian Center for Astrophysics. Since 1996, he has focused his research activity on the design and development of cryogenic LNAs within the LNA Laboratory, Observatorio Astronómico

Nacional (OAN), where he has carried out designs for European Space Agency (ESA), IRAM, Herschel, and ALMA among others.



**Carmen Díez-González** was born in Santander, Spain, in 1971. She received the Eng. degree in telecommunication in 1997.

From 1998 to 2000, she was involved with the characterization of sub-millimeter wave absorbers and far-infrared p-Ge lasers with SRON and with the Department of Applied Physics, Delft Technical University. From 2000 to 2004, she was with TTI, giving support to the Observatorio Astronómico Nacional (OAN) in the development of IF amplifiers for the European Space Agency (ESA)'s Herschel mission receivers. Since 2004, she has been involved with LNA Laboratory, Observatorio Astronómico Nacional (OAN), where she continues to be involved with Herschel activities, as well as with ALMA and other tasks of the laboratory.



**Inmaculada Malo-Gómez** was born in Madrid, Spain, in 1972. She received the Eng. degree in telecommunication in 1998 and the Doctor degree in telecommunication in 2011.

From 1996 to 2002, she was with telecommunications companies (Telefónica and Orange). In 2003, she joined the Observatorio Astronómico Nacional (OAN). Since then, she has been involved in the development and construction of low-noise cryogenic receivers for Yebes radio telescopes. Her main research interest is in passive and active components

for radio astronomy instrumentation.



**Enrique Villa** was born in Santander, Spain. He received the Telecommunications Engineering degree and Master's degree in information technologies and mobile networks communications from the University of Cantabria, Santander, Spain, in 2005 and 2008, respectively, and is currently working toward the Ph.D. degree at the University of Cantabria.

His research activity is related to the design and perform of radio-astronomy receivers, taking special interest in cryogenic behavior of phase switches, LNAs and detectors, and design, analysis. and

testing of microwave devices.



**Eduardo Artal** received the Engineer and Dr. Engineer in Telecommunication degrees from the Technical University of Catalonia, Barcelona, Spain, in 1976 and 1982, respectively.

From 1976 to 1990, he was an Assistant Professor with the Technical University of Catalonia. From 1979 to 1981, while on partial leave from the Technical University of Catalonia, he was with Mier Allende S.A., Barcelona, Spain, where he was involved with TV and FM radio re-emitters development. Since 1990, he has been a Professor

with the University of Cantabria, Santander, Spain, where he was Manager of their telecommunication engineering course from 1990 to 1994. From 1994 to 1998, he was Manager of the National Program for Information and Communications Technologies at the "Plan Nacional de I+D", National Research and Development Plan of the Spanish Ministry of Education and Science, Madrid, Spain. His main areas of activities and contributions have been microwave circuits and systems, including monolithic microwave integrated circuits up to 50 GHz. His current research interests are low-noise millimeter-wave amplifiers and receivers for radio-astronomy applications.